

# FastDIRC Status Report

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DIRC workfest  
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# Outline

- A proper (re-)introduction to FastDIRC
- Current status: reconstruction and observable studies
- Moving forward

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# The many meanings of “FastDIRC”

Depending on the context, the term “FastDIRC” can refer to:

- A reconstruction philosophy/approach
- A way of doing fast photon generation & propagation
- A specific implementation, i.e. a code repo on GitHub, of this reconstruction approach

Some charged track hits a DIRC bar

track's  
extrapolated  $\vec{p}, \vec{x}, t$

A “black box” that we try to model as best as we can: geometry, photon generation and propagation

O(50) observed PMT hits

each hit: PMT channel  
and time

# FastDIRC reconstruction philosophy

Some charged track hits a DIRC bar

track's  
extrapolated  $\vec{p}, \vec{x}, t$

A “black box” that we try to model as best as we can: geometry, photon generation and propagation

O(50) observed  
PMT hits

each hit: PMT channel  
and time

What if...

O(100k or 1M) hits that could have been generated by this track, according to our model of the “black box”

Let's call them “**support points**”

# FastDIRC reconstruction philosophy

track's  
extrapolated  $\vec{p}, \vec{x}, t$

$O(1M)$  support points  
under **pion** hypothesis

$O(1M)$  support points  
under **kaon** hypothesis

Given the  $O(50)$  observed hits and the  $O(1M)$  support points under each hypothesis, how likely do the observed hits come from each hypothesis?

# FastDIRC reconstruction philosophy: a 3D-PDF approach

Answer: Hypothesis testing.

Compute the likelihood under each hypothesis, and form delta log-likelihoods

How to compute the likelihood under each hypothesis?

Pseudocode\*:

likelihood = 1

**for** every observed hit  $O_i$  **do**

**for** every support point  $S_j$  **do**

$$r_{i,j}^2 = \frac{(x_{O_i} - x_{S_j})^2}{s_x^2} + \frac{(y_{O_i} - y_{S_j})^2}{s_y^2} + \frac{(t_{O_i} - t_{S_j})^2}{s_t^2}$$

$$\text{likelihood} *= \exp\left(-\frac{r_{i,j}^2}{s_b^2}\right)$$

**end for**

**end for**

\*In the real code, cut-offs, logs and additions are used in intermediate steps to make computation tractable.



# FastDIRC reconstruction philosophy: a 3D-PDF approach

$$r_{i,j}^2 = \frac{(x_{O_i} - x_{S_j})^2}{s_x^2} + \frac{(y_{O_i} - y_{S_j})^2}{s_y^2} + \frac{(t_{O_i} - t_{S_j})^2}{s_t^2}$$
$$\text{likelihood} \stackrel{*}{=} \exp\left(-\frac{r_{i,j}^2}{s_b^2}\right)$$

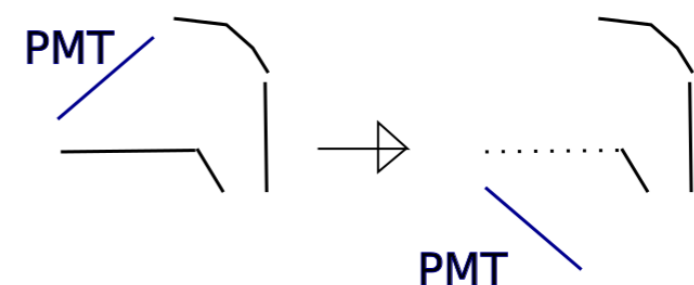
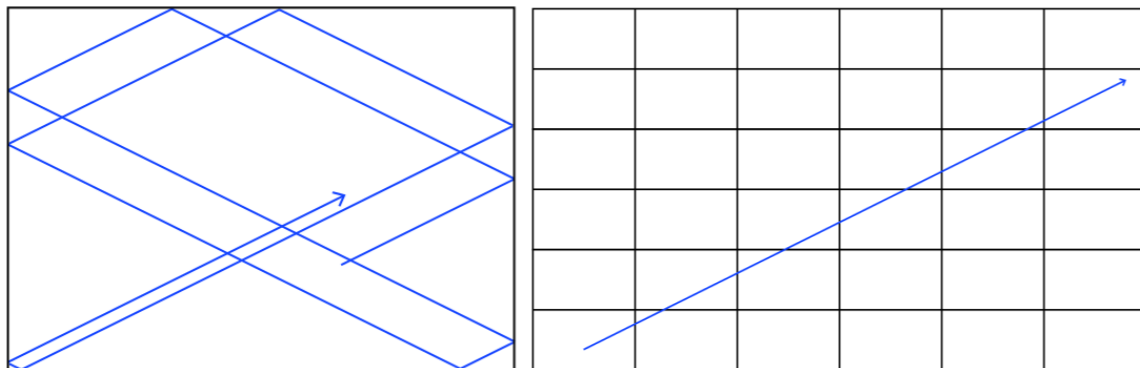
- $s_x, s_y, s_t, s_b$  are simply free parameters that represent some sort of scaling and they do *not* correspond to resolutions on those quantities
- This is simply *one* way of computing likelihoods from the given support points and observed hits, i.e. different kernels can be used
- At the present stage, it's likely more productive to focus on the question of *whether the support points represent the 3 dimensional probability density function from which the observed points are drawn*

How do you obtain the  $O(1M)$  support points per charged track?

- **Monte Carlo approach**, i.e. Geant
  - pros: reasonably model all the effects that you put in
  - cons: slow, cannot do this for every track — naive implementation would render this reconstruction not viable
- **Analytical approach**
  - pros: fast enough to make this reconstruction approach feasible
  - cons: cannot model everything — could miss subtle effects

# FastDIRC simulation: a fast analytical ray-tracing modeling

- Generation: only generate those that will be detected
  - Sample a wavelength spectrum: intrinsic  $1/\lambda^2$ , materials'  $\lambda$  dependence, PMT Q.E. etc.
  - Determine the index of refraction:  $n(\lambda)$
  - Throw the photons around the Cherenkov cone
- Propagation: purely geometrical
  - A reflection: e.g.  $\text{dirVec}_x = -\text{dirVec}_x$
  - Reflections are perfect
  - Gaussian smearing can be added



# FastDIRC simulation: a fast analytical ray-tracing modeling

- The modeling of generation+propagation is certainly not comprehensive
- It can serve as an independent way of checking data, e.g. photon yield, timing
- It *enables* the 3D-PDF reconstruction approach, but in principle it is also *independent* of the reconstruction approach used

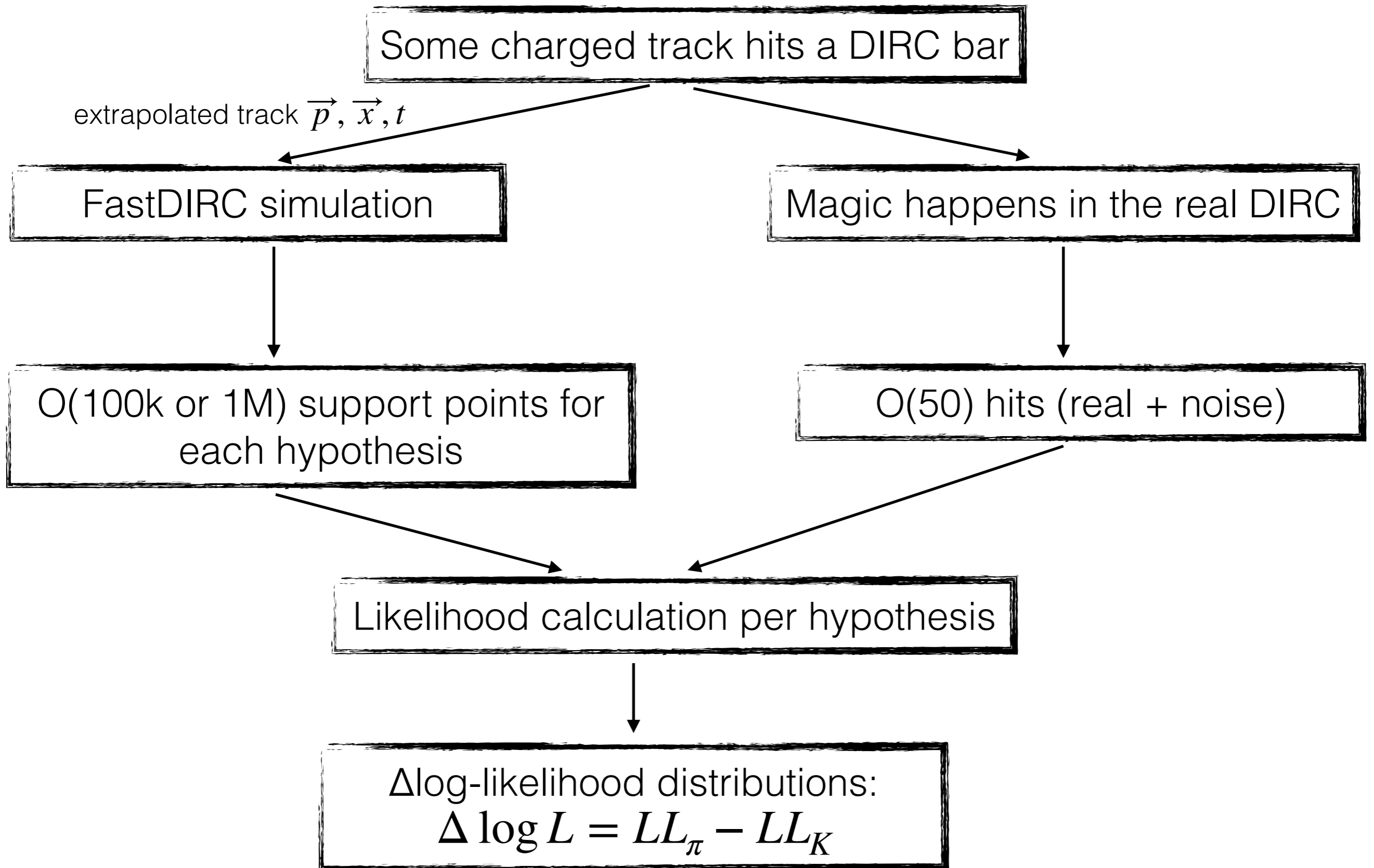
Key questions I am focusing on:

- The ability to quickly simulate a large number of potential PMT hits should be a powerful tool at our disposal. How can we make the best use of it?
- The correct modeling of the p.d.f that the observed hits are drawn from is critical. How can we ensure that?

# Outline

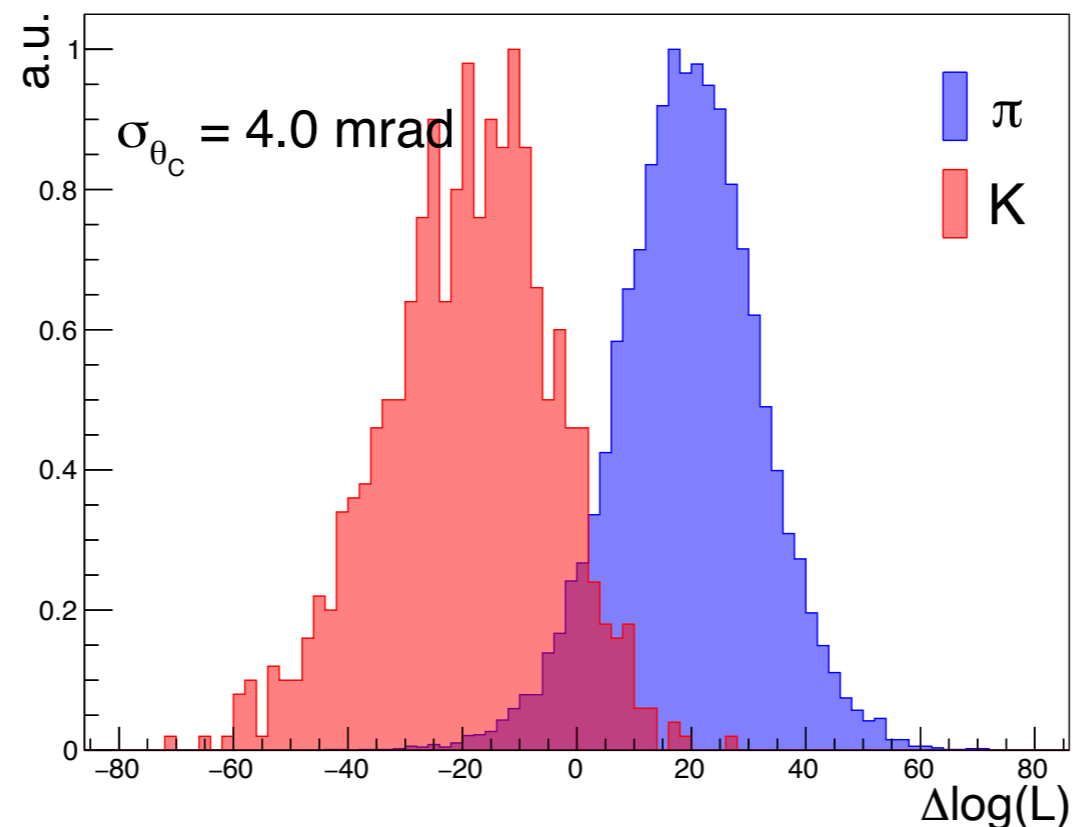
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# FastDIRC reconstruction



To convert from a  $\Delta LL$  distribution into an equivalent angular resolution:

- From  $\Delta LL$ , construct the ROC curve and compute the AUC
- In  $\theta_C$  space, we know the theoretical  $\Delta\theta_C$  for this momentum ( $\sim 11.5$  mrad in this case)
- Assume both the pion and kaon  $\theta_C$  are Gaussian-distributed and the single track resolution  $\sigma_{\theta_c}$  is the same for both, numerically compute what  $\sigma_{\theta_c}$  value would produce the same AUC



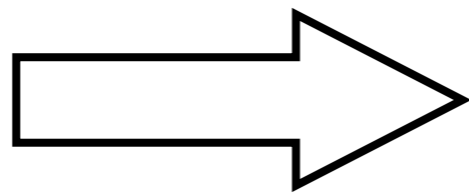
From a sample of 3 GeV  $\pi+K$  tracks

# FastDIRC current status

- Adapted the original FastDIRC code with the as-built nominal geometry
- Functional as an alternative reconstruction method
- Achieved  $\pi/K$  separation ( $\sim 3\sigma$  at 3 GeV) comparable with the without-any-correction geometrical reconstruction
- Experimented with some ideas:
  - Attempted at alignment with geometry parameters: looked at different figures-of-merit, tried a few naive overall shifts
  - Construction of low-level observables: as an independent check of data quality (during commissioning and beyond), and to identify where the inner workings of the modeling need to be improved

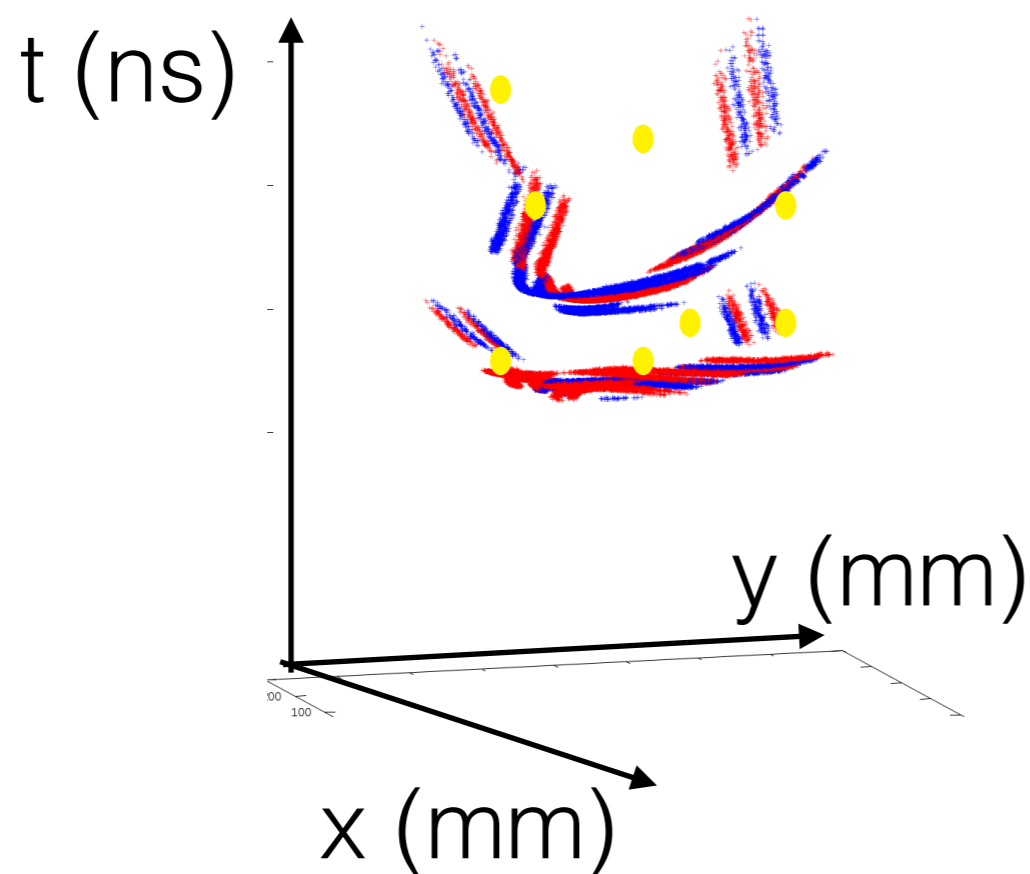


Input:  
track kinematics,  
mass hypothesis



$O(1M)$  support points  
in  $(x, y, t)$  space

Model:  
geometry + propagation



red and blue bands:  
support points from  
two mass hypos

● observed hit

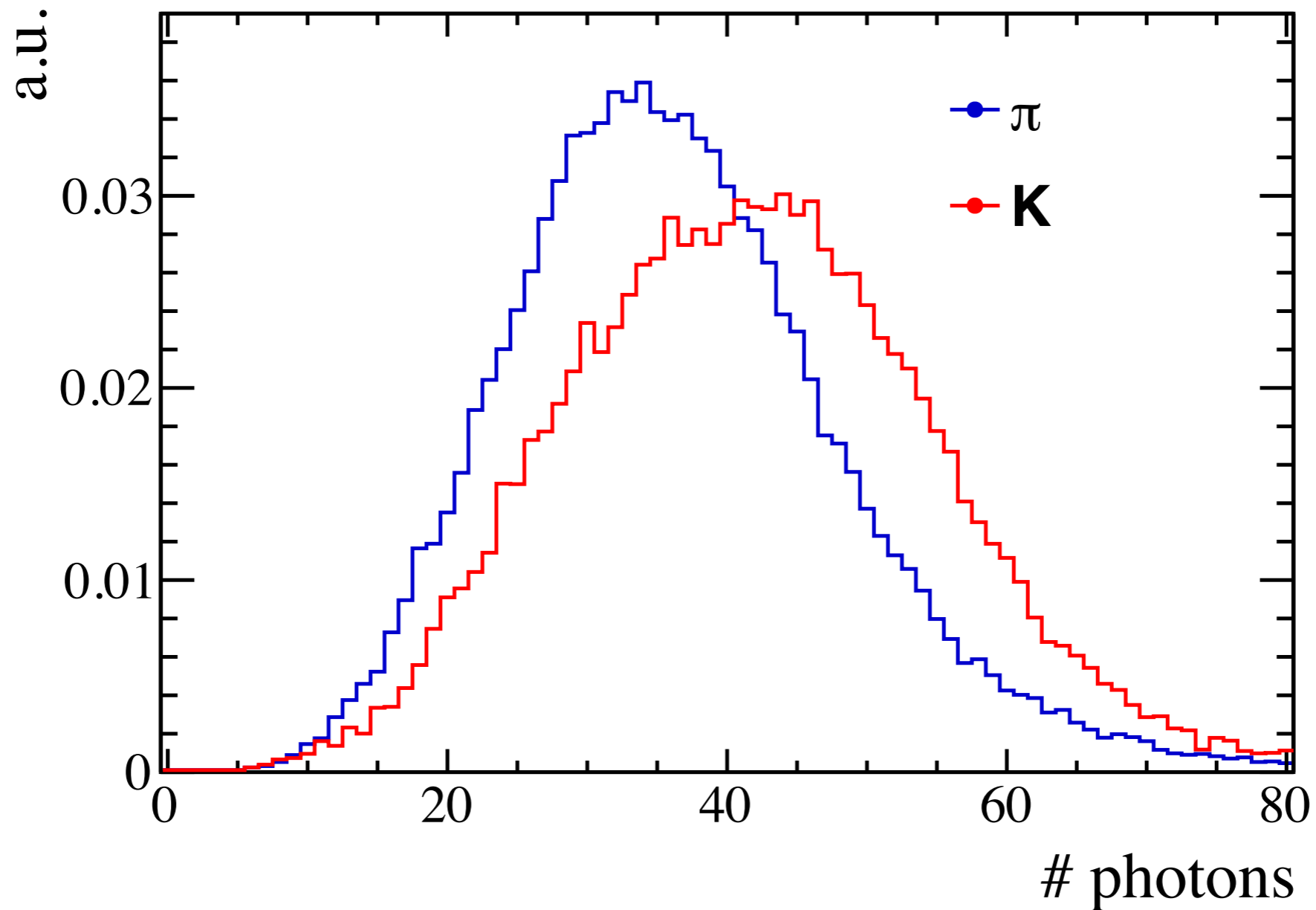
Idea: for each observed hit, look at its “neighbors” in this 3D space, and try to construct observables

For each hit, define its *neighborhood* as, e.g. within  $\pm 5$  ns and  $\pm 8.5$  mm:

- Yes/No counting  $\Leftrightarrow$  photon yield:
  - If there is at least one (no) support points in the neighborhood, call it a signal (noise) hit
- Distance measure  $\Leftrightarrow$  Delta observables:
  - For all support points in the neighborhood, construct  $\Delta x = x_0 - x_{S,i}$ ,  $\Delta y = y_0 - y_{S,i}$ ,  $\Delta t = t_0 - t_{S,i}$

$\Rightarrow$  Showed interesting/promising signs, but a lot to understand

# Example: Photon Yield

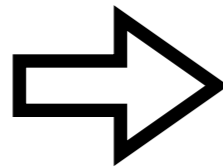
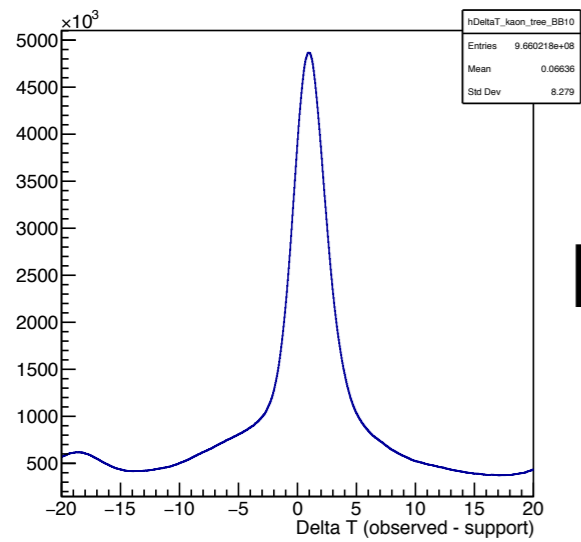


From a subset of 2019-11 dataset, integrated over 4+ GeV tracks and all locations on the DIRC

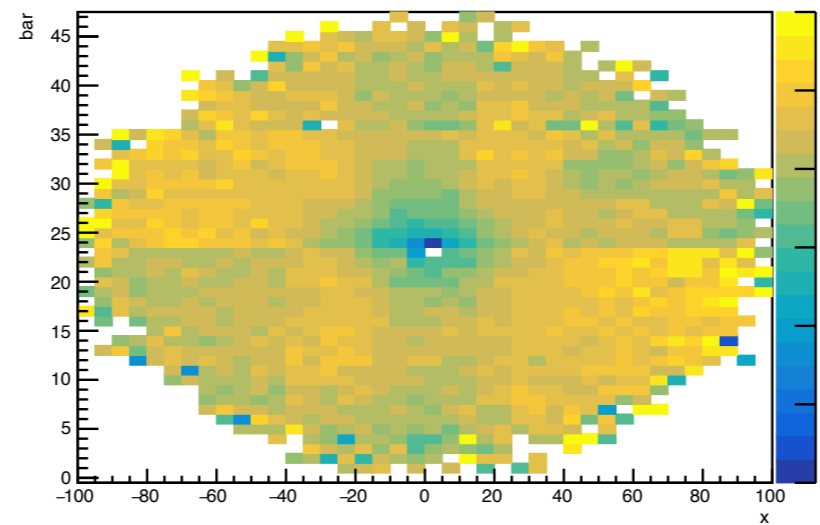
Similar behavior, but with an independent approach

# Example: $\Delta t$

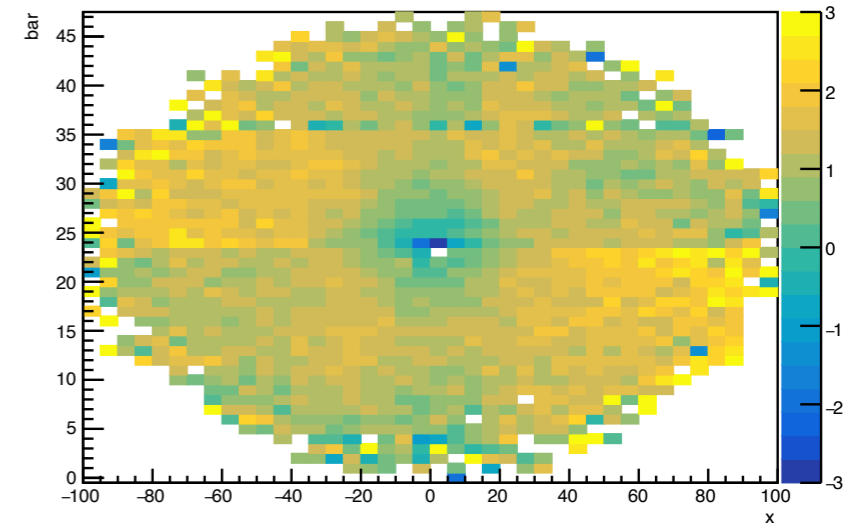
Per bar per x-bin



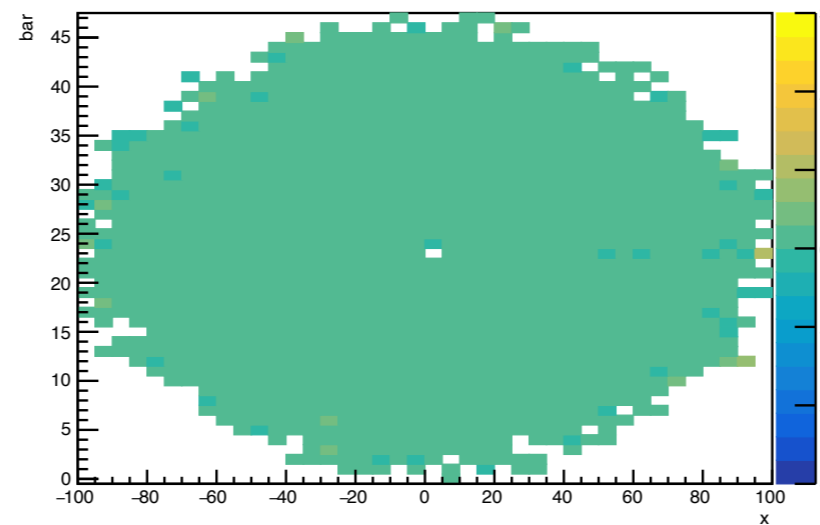
DeltaT  $\pi$  (data)



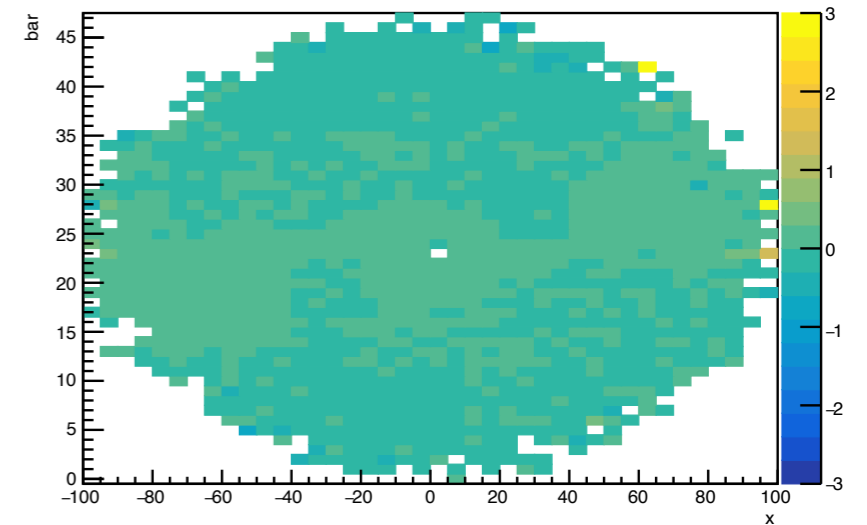
DeltaT K (data)



DeltaT  $\pi$  (sim hits)



DeltaT K (sim hits)



From a subset of 2019-11 dataset, 3 GeV tracks selected

Need to understand: what's the cause of overall shift?  
path length dependence?

This is just the beginning of looking at real data with this reconstruction approach. There is **A LOT** to understand:

- What does it mean to apply “corrections” in this context? (maybe this is not the right question to ask)
- How to factorize the various effects at play? What observables can we construct to do that?
- Fast simulation of support points should be a powerful tool, but how can we best use it?
- ...

Many aspects of the “code commissioning” still needed:

- Geometry manipulation works as intended
- Cherenkov wavelength spectrum represents our current knowledge

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# Future Plan

I need to graduate at some point... but before that, I hope to

- integrate FastDIRC into GlueX software
- continue exploring different ideas

# Summary

- FastDIRC is still useful and relevant in the era of data
- Its potential is not well unexplored
- It is just the beginning and I hope the story continues